# Markscheme 

May 2017

Physics

## Standard level

## Paper 3

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## Section A

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | a |  | in order to keep the temperature constant $\checkmark$ in order to allow the system to reach thermal equilibrium with the surroundings/OWTTE $\checkmark$ | Accept answers in terms of pressure or volume changes only if clearly related to reaching thermal equilibrium with the surroundings. | 1 max |
|  | b |  | ```recognizes b as gradient } calculates b}\mathrm{ in range 4.7 < 104 to 5.3 1 104 } Pam }``` | Award [2 max] if POT error in $b$. <br> Allow any correct SI unit, eg $\mathrm{kg} \mathrm{s}^{-2}$. | 3 |
|  | C |  | $V \propto H$ thus ideal gas law gives $p \propto \frac{1}{H} \checkmark$ <br> so graph should be «a straight line through origin,» as observed $\checkmark$ |  | 2 |
|  | d |  | $\begin{aligned} & n=\frac{b A}{R T} O R \text { correct substitution of one point from the graph } \checkmark \\ & n=\frac{5 \times 10^{4} \times 1.3 \times 10^{-3}}{8.31 \times 300}=0.026 \approx 0.03 \end{aligned}$ | Answer must be to 1 or 2 SF. Allow ECF from (b). | 2 |


| Question |  | Answers | Total |
| :--- | :--- | :--- | :--- | :--- | :--- |
| $\mathbf{1}$ | $\mathbf{e}$ | very large $\frac{1}{H}$ means very small volumes / very high pressures $\checkmark$ <br> at very small volumes the ideal gas does not apply <br> OR <br> at very small volumes some of the assumptions of the kinetic theory of gases do not hold $\checkmark$ |  |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 2 | a | $\begin{aligned} & g=\frac{4 \pi^{2} \times 1.60}{2.540^{2}}=9.7907 \\ & \Delta g=g\left(\frac{\Delta L}{L}+2 \times \frac{\Delta T}{T}\right)=« 9.7907 \times\left(\frac{0.01}{1.60}+2 \times \frac{0.005}{2.540}\right)=» 0.0997 \\ & \text { OR } \\ & 1.0 \% \checkmark \\ & \text { hence } g=(9.8 \pm 0.1) « \mathrm{~ms}^{-2} » \text { OR } \Delta g=0.1 « \mathrm{~ms}^{-2} » \checkmark \end{aligned}$ | For the first marking point answer must be given to at least 2 dp . <br> Accept calculations based on $\begin{aligned} & g_{\max }=9.8908 \\ & g_{\min }=9.6913 \\ & \frac{g_{\max }-g_{\min }}{2}=0.099 \approx 0.1 \end{aligned}$ | 3 |
|  | b | $\begin{aligned} & \frac{T}{T_{0}}=1.01 \checkmark \\ & \theta_{\max }=22 «^{\circ} » \checkmark \end{aligned}$ | Accept answer from interval 20 to 24. | 2 |

## Section B

Option A - Relativity

| Question |  | Answers | Notes | Total |
| :---: | :---: | :--- | :--- | :--- | :---: |
| $\mathbf{3}$ | a | a set of coordinate axes and clocks used to measure the position «in space/time of an object <br> at a particular time» <br> OR <br> a coordinate system to measure x,y,z,and $t /$ OWTTE $\checkmark$ | $\mathbf{1}$ |  |
|  | b | i | magnetic only $\checkmark$ <br> there is a current but no «net» charge «in the wire» $\checkmark$ |  |
|  | b | ii | electric only $\checkmark$ <br> $P$ is stationary so experiences no magnetic force $\checkmark$ <br> relativistic contraction will increase the density of protons in the wire $\checkmark$ | $\mathbf{2}$ |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 4 | a |  | $\Delta t_{\text {p }} /$ observer sitting in the train $\checkmark$ |  | 1 |
|  | b |  | $\gamma=\frac{\Delta t_{Q}}{\Delta t_{P}}=«=\frac{1}{0.30} »=3.3 \checkmark$ <br> to give $v=0.95 \mathrm{c} \checkmark$ |  | 2 |
|  | C |  | $\gamma=1.25 \checkmark$ <br> «length of train according $Q$ » $=125 / 1.25 \checkmark$ "giving 100 m " |  | 2 |
|  | d | i |  <br> axes drawn with correct gradients of $\frac{5}{3}$ for $c t^{\prime}$ and 0.6 for $x^{\prime} \checkmark$ | Award [1] for one gradient correct and another approximately correct. | 1 |


| $\mathbf{4}$ | d | ii |  |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- |

## Option B — Engineering physics

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 5 | a | i | $\frac{M}{3} v R \checkmark$ |  | 1 |
|  | a | ii | evidence of use of: $L=I \omega=\left(M R^{2}+\frac{M}{3} R^{2}\right) \omega \checkmark$ |  | 1 |
|  | a | iii | evidence of use of conservation of angular momentum, $\frac{M v R}{3}=\frac{4}{3} M R^{2} \omega \checkmark$ «rearranging to get $\omega=\frac{v}{4 R}$ " |  | 1 |
|  | a | iv | initial $K E=\frac{M v^{2}}{6} \checkmark$ <br> final $K E=\frac{M v^{2}}{24} \checkmark$ <br> energy loss $=\frac{M v^{2}}{8} \checkmark$ |  | 3 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 5 | b | i | $\begin{aligned} & \alpha «=\frac{3}{4} \frac{\Gamma}{M R^{2}} »=\frac{3}{4} \frac{0.01}{0.7 \times 0.5^{2}} \\ & \text { «to give } \alpha=0.04286 \text { rads }^{-2} » \end{aligned}$ | Working OR answer to at least 3 SF must be shown | 1 |
|  | b | ii | $\begin{aligned} & \theta=\frac{\omega_{i}^{2}}{2 \alpha} « \text { from } \omega_{f}^{2}=\omega_{i}^{2}+2 \alpha \theta » \checkmark \\ & \theta «=\frac{v^{2}}{32 R^{2} \alpha}=\frac{2.1^{2}}{32 \times 0.5^{2} \times 0.043} »=12.8 \text { OR } 12.9 \text { «rad» } \\ & \text { number of rotations «= } \frac{12.9}{2 \pi} »=2.0 \text { revolutions } \checkmark \end{aligned}$ |  | 3 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 6 | a |  | «a process in which there is» no thermal energy transferred between the system and the surroundings $\checkmark$ |  | 1 |
|  | b |  | A to B AND C to D $\checkmark$ |  | 1 |
|  | c | i | $\begin{aligned} & T=\frac{P V}{n R} \checkmark \\ & T\left(=\frac{512 \times 10^{3} \times 1.20 \times 10^{-3}}{0.150 \times 8.31}\right) \approx 493 « \mathrm{~K} » \end{aligned}$ | The first mark is for rearranging. | 2 |
|  | c | ii | $\begin{aligned} & P_{B}=\frac{P_{\mathrm{a}} V_{A}}{V_{B}} \checkmark \\ & P_{B}=267 \mathrm{kPa} \checkmark \end{aligned}$ | The first mark is for rearranging. | 2 |
|  | d | i | «B to C adiabatic so» $P_{B} V_{B}^{\frac{5}{3}}=P_{C} V_{C}^{\frac{5}{3}}$ AND $P_{C} V_{C}=n R T_{C}$ «combining to get result» $\checkmark$ | It is essential to see these 2 relations to award the mark. | 1 |
|  | d | ii | $\begin{aligned} & T_{C}=\left(\frac{P_{B} V_{B}^{\frac{5}{3}}}{n R}\right) V_{C}^{\frac{-2}{3}} \checkmark \\ & T_{C}=«\left(\frac{267 \times 10^{3} \times\left(2.30 \times 10^{-3}\right)^{\frac{5}{3}}}{0.150 \times 8.31}\right)\left(2.90 \times 10^{-3}\right)^{\frac{-2}{3}} 》=422 « \mathrm{~K} » \checkmark \end{aligned}$ |  | 2 |
|  | e |  | the isothermal processes would have to be conducted very slowly / OWTTE $\checkmark$ |  | 1 |

## Option C - Imaging

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 7 | a | i | an image formed by extensions of rays, not rays themselves OR <br> an image that cannot be projected on a screen $\checkmark$ |  | 1 |
|  | a | ii | $\begin{aligned} & \frac{1}{v}=\frac{1}{3.0}-\frac{1}{4.0} \checkmark \\ & « v=12 \mathrm{~cm} » \end{aligned}$ |  | 1 |
|  | a | iii | $\begin{aligned} & u=18-12=6.0 « \mathrm{~cm} » \\ & v=-24 \text { «cm» } \downarrow \\ & « \frac{1}{f}=\frac{1}{6.0}-\frac{1}{24} \Rightarrow>f=8.0 « \mathrm{~cm} » \end{aligned}$ | Award [2 max] for answer of 4.8 cm . <br> Minus sign required for MP2. | 3 |
|  | a | iv | line parallel to principal axis from intermediate image meeting eyepiece lens at $P$ line from arrow of final image to $P$ intersecting principal axis at $F \checkmark$ |  | 2 |


| Question |  | Answers | Notes | Total |  |
| :--- | :--- | :--- | :--- | :--- | :---: |
| $\mathbf{7}$ | b | i | object is far away so intermediate image forms at focal plane of objective $\checkmark$ <br> for final image at infinity object must also be at focal point of eyepiece $\checkmark$ <br> «hence 87.5 cm» | No mark for simple addition of <br> focal lengths without <br> explanation. |  |
|  | b | ii | angular magnification $=\frac{85.0}{2.50}=34 \checkmark$ <br> angular diameter $34 \times 7.8 \times 10^{-3}=0.2652 \approx 0.27$ «rad» $\checkmark$ |  | 2 |
| c | chromatic aberration is the dependence of refractive index on wavelength $\checkmark$ <br> but mirrors rely on reflection <br> OR <br> mirrors do not involve refraction $\checkmark$ <br> «so do not suffer chromatic aberration» | 2 |  |  |  |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 8 | a | i | Ionger distance without amplification signal cannot easily be interfered with $\checkmark$ less noise $\checkmark$ <br> no cross talk $\checkmark$ <br> higher data transfer rate $\checkmark$ |  | 2 max |
|  | a | ii | infrared radiation suffers lower attenuation $\checkmark$ |  | 1 |
|  | b |  | $\begin{aligned} & \text { loss }=10 \log \frac{2.4}{15} «=-7.959 \mathrm{~dB} » \\ & \text { length }=« \frac{7.959}{0.30}=» 26.53 \approx 27 \text { «km» } \end{aligned}$ |  | 2 |
|  | C |  | a thin core means that rays follow essentially the same path / OWTTE $\checkmark$ and so waveguide (modal) dispersion is minimal / OWTTE $\checkmark$ |  | 2 |

## Option D - Astrophysics

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 9 | a | i | stars fusing hydrogen «into helium» $\checkmark$ |  | 1 |
|  | a | ii | $\begin{aligned} & M=M_{\odot}\left(4 \times 10^{5}\right)^{\frac{1}{3.5}}=39.86 M_{\odot} \checkmark \\ & « M \approx 40 M_{\odot}> \end{aligned}$ | Accept reverse working. | 1 |
|  | a | iii | $\begin{aligned} & 4 \times 10^{5}=13^{2} \times \frac{T^{4}}{6000^{4}} \downarrow \\ & T \approx 42000 « \mathrm{~K} » \checkmark \end{aligned}$ | Accept use of substituted values into $L=\sigma 4 \pi R^{2} T^{4}$. <br> Award [2] for a bald correct answer. | 2 |
|  | a | iv | $\begin{aligned} & 4 \times 10^{-11}=4 \times 10^{5} \times \frac{1 \mathrm{AU}^{2}}{d^{2}} \\ & d=1 \times 10^{8} \text { «AU» } \end{aligned}$ | Accept use of correct values into $b=\frac{L}{4 \pi d^{2}}$. | 2 |
|  | b |  | the gravitation «pressure» is balanced by radiation «pressure» $\checkmark$ that is created by the production of energy due to fusion in the core / OWTTE $\checkmark$ | Award [1 max] if pressure and force is inappropriately mixed in the answer. <br> Award [1 max] for unexplained "hydrostatic equilibrium is reached". | 2 |


| Question |  | Answers | Total |  |
| :--- | :--- | :--- | :--- | :---: | :---: |
| $\mathbf{9}$ | $\mathbf{c}$ | the Sun will evolve to become a red giant whereas Theta 1 Orionis will become a red super <br> giant $\checkmark$ <br> the Sun will explode as a planetary nebula whereas Theta 1 Orionis will explode as a <br> supernova $\checkmark$ <br> the Sun will end up as a white dwarf whereas Theta 1 Orionis as a neutron star/black hole $\checkmark$ |  |  |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 10 | a |  | black body radiation / $3 \mathrm{~K} \checkmark$ <br> highly isotropic / uniform throughout <br> OR <br> filling the universe $\checkmark$ | Do not accept: CMB provides evidence for the Big Bang model. | 2 |
|  | a | ii | $« \lambda=\frac{2.9 \times 10^{-3}}{2.8} » \approx 1.0<\mathrm{mm} » \checkmark$ |  | 1 |
|  | b |  | the universe is expanding and so the wavelength of the CMB in the past was much smaller $\checkmark$ indicating a very high temperature at the beginning $\checkmark$ |  | 2 |
|  | c | i | $\begin{aligned} & « z=\frac{v}{c} \Rightarrow » v=0.16 \times 3 \times 10^{5} «=0.48 \times 10^{5} \mathrm{~km} \mathrm{~s}^{-1} » \checkmark \\ & « d=\frac{v}{H_{0}} \Rightarrow v=\frac{0.48 \times 10^{5}}{68}=706 » \approx 710 « \mathrm{Mpc} » \end{aligned}$ | Award [1 max] for POT error. | 2 |
|  | c | ii | $\begin{aligned} & z=\frac{R}{R_{0}}-1 \Rightarrow \frac{R}{R_{0}}=1.16 \checkmark \\ & \frac{R_{0}}{R}=0.86 \checkmark \end{aligned}$ |  | 2 |

